## ANALYSIS OF A AND E COMPONENTS OF THE $\nu_{21}$ TORSIONAL FUNDAMENTAL OF PROPENE AT 188 cm<sup>-1</sup>

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The far infrared spectrum of propene,  $CH_3 - CH = CH_2$ , has been measured at 193.65 K with a resolution of 0.00096 cm<sup>-1</sup> on the FIR beamline at the Canadian Light Source (CLS) synchrotron. The torsional fundamental band  $\nu_{21}$  at 188 cm<sup>-1</sup> is quite complex because the coupling of the angular momenta of the overall rotation with the internal rotation of the CH<sub>3</sub> group splits every rovibrational energy level (without internal rotation) into an *A*- and an *E*-component. As a result, every rovibrational transition splits into two or three components. To predict and assign the spectrum and to determine spectroscopic constants by the LS method, the ERHAM program <sup>*b*</sup> has been modified to use it for rovibrational spectra.

Initially, the spectroscopic constants for the ground state (GS) were fixed at the values determined by Craig et al.<sup>c</sup> from almost 900 microwave and millimeter-wave frequencies. This was good enough to begin the assignment of transitions of  $\nu_{21}$  involving low  $K_a$  levels. However, for  $K_a > 3$ , the ratio of the standard deviation over the estimated experimental uncertainty increased with increasing  $K_a$  despite the introduction of additional tunneling parameters.

Only the inclusion of the internal rotation parameters  $\rho$  and  $\beta$  (which had been held constant with the GS constants) among the variable parameters brought the standard deviation down to an acceptable level. From then on, the transitions of  $\nu_{21}$  were used simultaneously with all GS transitions (footnote b) to fit the parameters of two different effective internal rotation Hamiltonians (one each for the GS and  $\nu_{21}$  levels) but with shared  $\rho$  and  $\beta$ . More than 6000 individual transitions (*J* up to 50,  $K_a$  up to 8) have been assigned so far for the  $\nu_{21}$  band and preliminary constants have been obtained, among them an improved internal rotation barrier.

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<sup>&</sup>lt;sup>b</sup>P. Groner, J. Chem Phys. (1997) 107, 4483; J. Mol. Spectrosc. (2012) 278, 52

<sup>&</sup>lt;sup>c</sup>N. C. Craig et al., J. Mol. Spectrosc. (2016) 328, 1