

## QUANTUM CHEMICAL MODELING OF ASTROCHEMICAL REACTIONS OF C ATOM AND C<sup>+</sup> CATION WITH NH<sub>3</sub> BOUND TO AMORPHOUS WATER ICE

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Reactions of C atoms and C<sup>+</sup> cations with NH<sub>3</sub> on water ice were characterized with density function theory using modestly sized water clusters. Both reactions are expected to occur in dense interstellar clouds and in protostellar sources. The neutral C(<sup>3</sup>P) + NH<sub>3</sub> reaction on ice begins with the formation of triplet CNH<sub>3</sub> via dative bonding involving the 2s<sup>2</sup> lone pair on nitrogen. Assuming it is not ejected into the gas phase, CNH<sub>3</sub> can subsequently react with one or two H atoms to yield CH<sub>2</sub>NH<sub>2</sub> and then CH<sub>3</sub>NH<sub>2</sub>. The ion-molecule C<sup>+</sup>(<sup>2</sup>P) + NH<sub>3</sub> reaction on ice begins with charge transfer so that C(<sup>3</sup>P) reacts with NH<sub>3</sub><sup>+</sup>. The short-lived CNH<sub>3</sub><sup>+</sup> intermediate, which has a covalent C-N bond, deprotonates to yield the H<sub>2</sub>NC radical, which was detected in 2021 toward the dark cloud L483 and other sources. Doublet H<sub>2</sub>NC can react with H atoms to yield several different products. The vibrational spectrum of NH<sub>3</sub> on amorphous ice will also be presented.