

## INVESTIGATION OF THE ZEEMAN EFFECT IN THE $e^6\Pi - a^6\Delta$ SYSTEM OF FeH: APPLICATION TO STELLAR SPECTROSCOPY

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We investigate the Zeeman effect in the  $e^6\Pi - a^6\Delta$  system of FeH. The molecule has been produced by reaction of  $\text{Fe}(\text{CO})_5$  with H atoms, and examined by laser excitation with selective detection and by Fourier transform resolved fluorescence. The field-free spectrum of the  $e^6\Pi - a^6\Delta$  system was extensively studied by the J.M. Brown group and collaborators in the 1990's<sup>a</sup>. Their analysis located the low-lying excited  $a^6\Delta$  state  $1890\text{ cm}^{-1}$  above the  $X^4\Delta$  electronic ground state.

One of the infrared systems of FeH already recognised<sup>b</sup> in the spectra of cool stars around  $1.6\ \mu\text{m}$ ,  $E^4\Pi - A^4\Pi$ , originates in an excited state lying only  $920\text{ cm}^{-1}$  lower in energy than  $a^6\Delta$ , so it seemed not unreasonable to expect signatures from the known sextet systems  $e-a$  and  $g-a$  to appear in the spectra of cool stars as well. We found that the  $e-a$  system can indeed be used as a diagnostic for FeH. Cross-correlation functions between a mask of laboratory-measured  $e-a$  transitions and spectra from four M-dwarf stars taken on spectropolarimetric instruments ESPaDOnS (on Maunakea) and Narval (in the French Pyrenees), show that reliable radial-velocity information can be extracted for these objects. Our new Zeeman measurements are intended to improve reference data for cross-correlation calculations for M-type stars, whose magnetic fields are typically  $0 - 5000$  Gauss.

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<sup>a</sup>Goodridge *et al* J. Chem. Phys. 106 (12), 4823 (1997); Wilson *et al* J. Chem. Phys. 115 (13), 5943 (2001)

<sup>b</sup>Wallace & Hinkle, ApJ, 559, 424 (2001)