

MAGNETIC FIELD STRENGTH LIMITS IN A PROTOPLANETARY DISK FROM MULTI-WAVELENGTH ZEEMAN OBSERVATIONS

RACHEL E. HARRISON, Astronomy, University of Illinois at Urbana-Champaign, Urbana, IL, USA; LESLIE LOONEY, AASSIK PAZHANI, Department of Astronomy, University of Illinois at Urbana-Champaign, Urbana, IL, USA; ZHI-YUN LI, Department of Astronomy, The University of Virginia, Charlottesville, VA, USA; HAIFENG YANG, Institute for Advanced Study, Tsinghua University, Beijing, China; IAN STEPHENS, Department of Earth, Environment, and Physics, Worcester State University, Worcester, MA, USA; RICHARD TEAGUE, Radio and Geoastronomy Division, Harvard-Smithsonian Center for Astrophysics, Cambridge, MA, USA; RICHARD CRUTCHER, Department of Astronomy, University of Illinois at Urbana-Champaign, Urbana, IL, USA; CRYSTAL L. BROGAN, NAASC, National Radio Astronomy Observatory, Charlottesville, VA, USA; ERIN GUILFOIL COX, Center for Interdisciplinary Exploration and Research in Astrophysics (CIERA), Northwestern University, Evanston, IL, USA.

Magnetic fields likely play a critical role in the accretion of material from protoplanetary disks onto protostars by providing a mechanism of angular momentum transport, particularly through magnetic disk winds. Constraining magnetic field strengths in protoplanetary disks is therefore necessary to test theories of magnetically-driven accretion. Zeeman splitting observations offer a way to directly measure or set upper limits on magnetic field strengths. We present the results of Zeeman splitting observations of several hyperfine lines of the CN(2-1) and CN(1-0) transitions in the Class II protoplanetary disk V4046 Sgr. We also present observations of the linear continuum dust polarization in this source and discuss their implications for the disk's dust population.